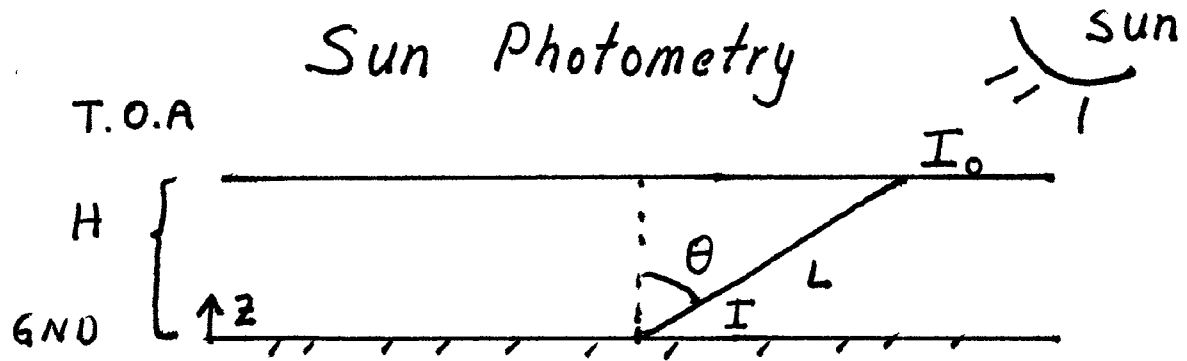




Let

H = "height" of atmosphere

L = Pathlength of sunlight through atmosphere.

θ = Solar Zenith Angle

Then $\cos \theta = \frac{H}{L}$, $L = \frac{H}{\cos \theta}$

Let

$m = 1/\cos \theta$ = Air Mass, $m \geq 1$

$m = 1$ Sun overhead

$m = 10$ Sunrise or sunset

$I_0 = I_0(\lambda)$ = T.O.A Solar irradiance

$I(\lambda)$ = GND solar irradiance

τ = optical depth - dimensionless quantity
(when the sun is overhead)

Then

Direct Beam

(no multiple scattering)

$$I(\lambda) = e^{-\frac{\tau(\lambda)}{\cos \theta}} I_0(\lambda)$$

Sunlight at surface

Sunlight at top of atmosphere

$$\text{So } I(\lambda) = e^{-m\tau(\lambda)} I_0(\lambda)$$

Idea: 1. Measure $I(\lambda)$ at the ground at different times, so with different m values.

2. The sun output $I_0(\lambda)$ at the top of the atmosphere remains the same.

3. Unknown: $\tau(\lambda)$ Solve for it, learn about the atmosphere with it.

Here's how:

$$\tau(\lambda) = \tau_{\text{gas}}(\lambda) + \tau_{\text{aerosol}}(\lambda) + \tau_{\text{cloud}}(\lambda)$$

$$\tau_{\text{gas}}(\lambda) \Rightarrow \tau_{\text{gas}}^{\text{scattering}}(\lambda) + \tau_{\text{gas}}^{\text{absorption}}(\lambda)$$

Rayleigh Scattering
(Blue Sky)

$\text{O}_2, \text{N}_2,$

$$= \frac{a}{\lambda^4}$$

Absorption by gases
like water vapor, O_2 ,
Ozone (O_3),
Nitrogen dioxide (NO_2)

$$\tau_{\text{aerosol}}(\lambda) = \tau_{\text{aerosol}}^{\text{scattering}}(\lambda) + \tau_{\text{aerosol}}^{\text{absorption}}(\lambda)$$

Sulfates

Nitrates

Sea Salt

Organic Carbon

Dust

Black Carbon

Black Carbon

(Soot)

$\sim \frac{c}{\lambda}$ Dust

(UV wavelengths)

$$\sim \frac{b}{\lambda^2}$$

Example: Rayleigh Scattering by O_2, N_2

$$\frac{\tau_{\text{gas}}^{\text{scattering}}(\lambda = 400\text{nm})}{\tau_{\text{gas}}^{\text{scattering}}(\lambda = 600\text{nm})} = \left(\frac{600}{400}\right)^4 = \frac{81}{16} \approx 5$$

Blue
Red

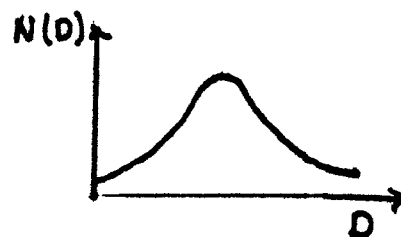
Conclusion: Scattering in blue by gases is
5 x scattering in Red.

Aerosol: $\tau_{\text{aerosol}}^{\text{ext}}(\lambda) = \int_0^H \int_0^\infty \sigma_{\text{ext}}(\lambda, D) N(D, z) dD dz$

ext = abs + sca

Particle extinction cross section as a function
of wavelength and particle size.

$N(D, z)$ = Particle Size Distribution
(D = particle diameter)



Gas Absorption:

$$\tau_{\text{gas}}^{\text{absorption}}(\lambda) = \int_0^H C(z) \sigma_{\text{abs}}(z, \lambda) dz$$

of
Absorbing Molecules
Volume

Absorption Cross
Section per Molecule

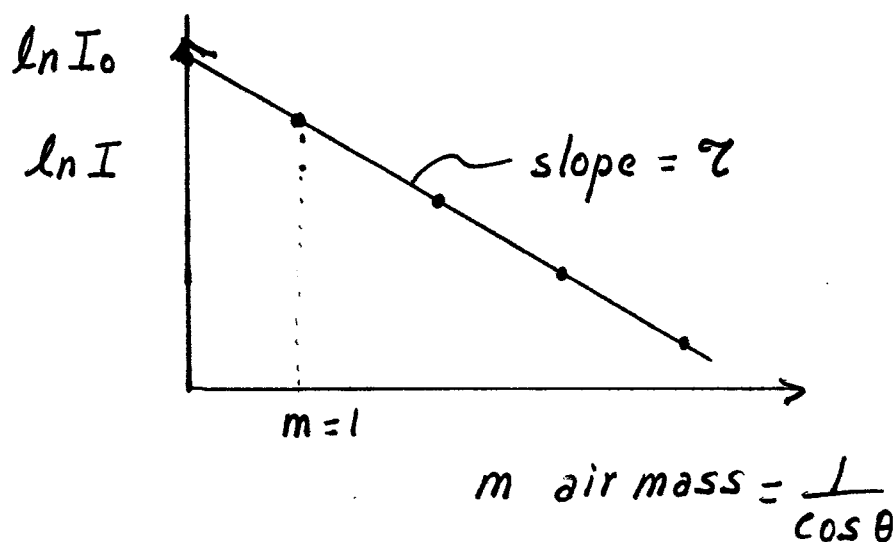
Here's how to crunch the data and get information from it.

Date	Time	Lat	Long	Calculate θ	Measure $I(\lambda)$ Volts	Measure $I(\lambda)$ Dark Volts
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$$I(\lambda) = e^{-\tau m} I_0(\lambda)$$

Then $\ln I = \ln I_0 - m\tau$

$\nearrow$ offset (fixed, sunlight at top of Atmosphere)
 $\uparrow$ independent variable
 $\searrow$ slope use to learn about atmosphere



How to process

$$\text{Slope} = \tau = \tau_{\text{gas}}^{\text{sca}} + \tau_{\text{aerosol}} + \tau_{\text{cloud}}$$

look up from table $\tau_{\text{gas}}^{\text{sca}}$ } aerosol amount (optical depth) τ_{aerosol} "Clear" day (No clouds) τ_{cloud}
 $\tau_{\text{aerosol}} \sim 0.2$ or so.

For 940nm, also get water vapor concentration

$$\tau = \tau_{\text{gas}}^{\text{scatt.}} + \tau_{\text{H}_2\text{O}}^{\text{absorp.}} + \tau_{\text{aerosol}}$$

look up $\tau_{\text{gas}}^{\text{scatt.}}$ } Large when atmosphere is moist $\tau_{\text{H}_2\text{O}}^{\text{absorp.}}$ Get from 870nm τ_{aerosol}

Wavelength (nm)

Comment

430		Good for aerosol, Rayleigh
470		scattering larger
530		
660		Good for aerosol, Rayleigh modest
870	—	Good for aerosol, Rayleigh small
940	—	water vapor channel

Also, get $I_0(\lambda)$ from the offset.

$$I_0(\lambda) = \overset{\text{calibrated}}{I_0(\lambda)} = K(\lambda) \overset{\text{counts or voltage}}{I_0(\lambda)} \quad \text{from offset}$$

↑
calibration factor

From satellite measurements, TABLE

$$K(\lambda) = \frac{\overset{\text{calibrated}}{I_0(\lambda)} \left[\frac{\text{Watts}}{\text{m}^2 \cdot \text{nm}} \right]}{\overset{\text{counts}}{I_0(\lambda)}}$$

Raw output from spectrometer.

Use $K(\lambda)$ to calibrate any spectrum from the spectrometer.

